

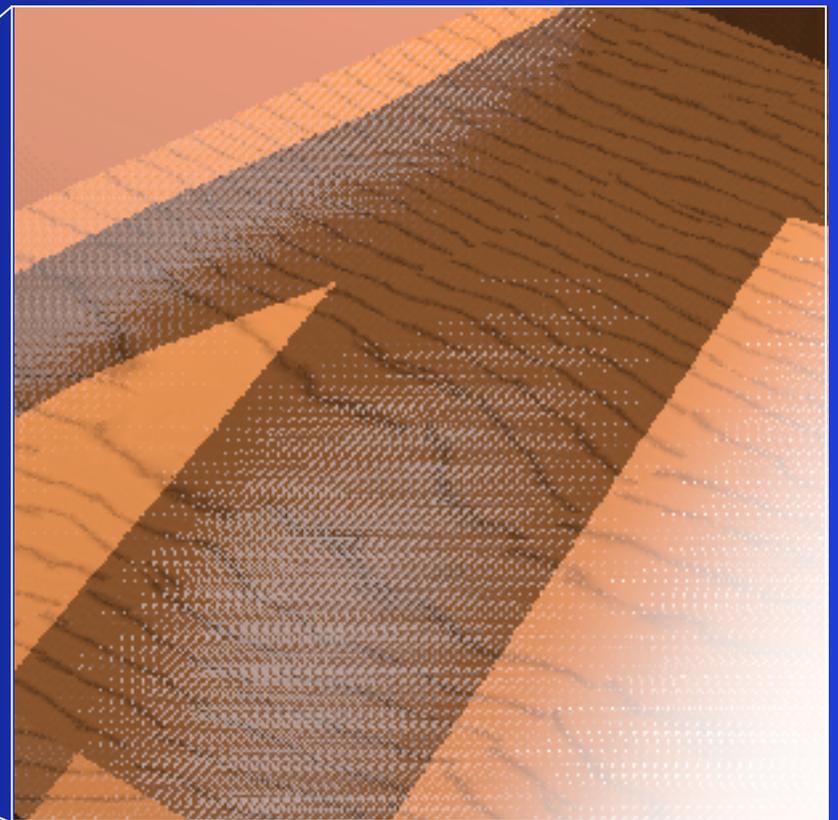
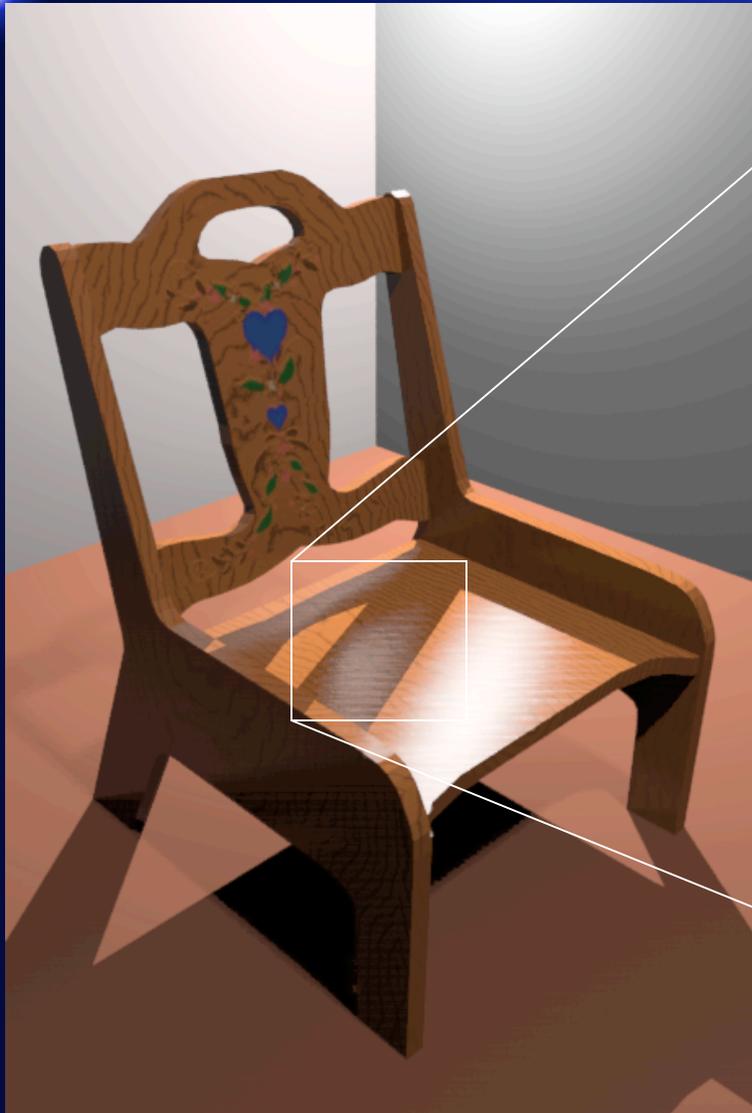
Code Changes and Additions in *Radiance* 3.7

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Anywhere Software

What's New in 3.7

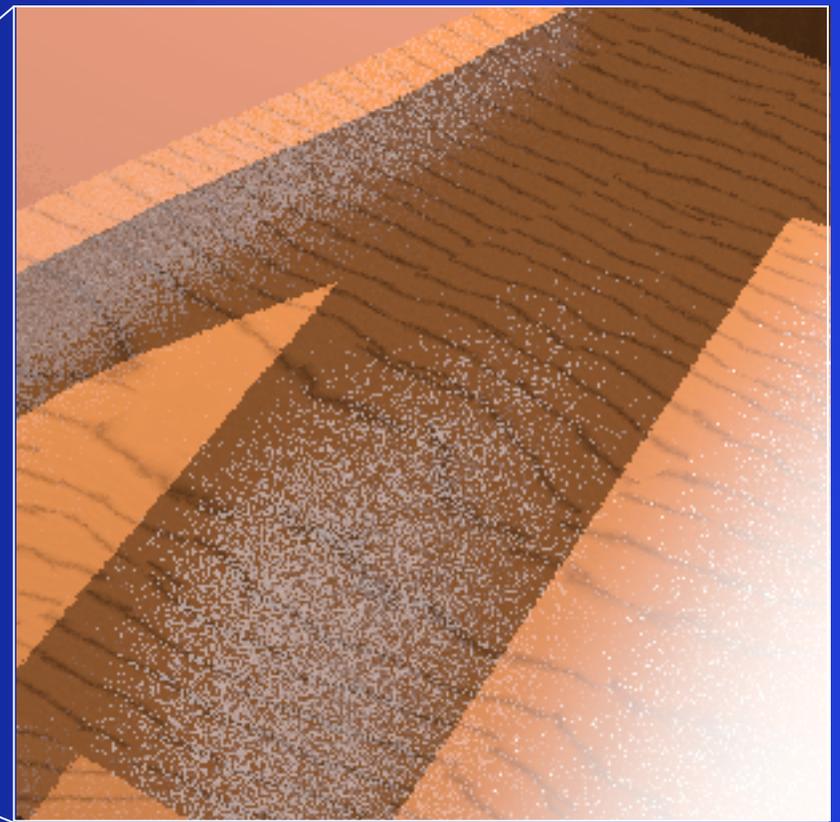
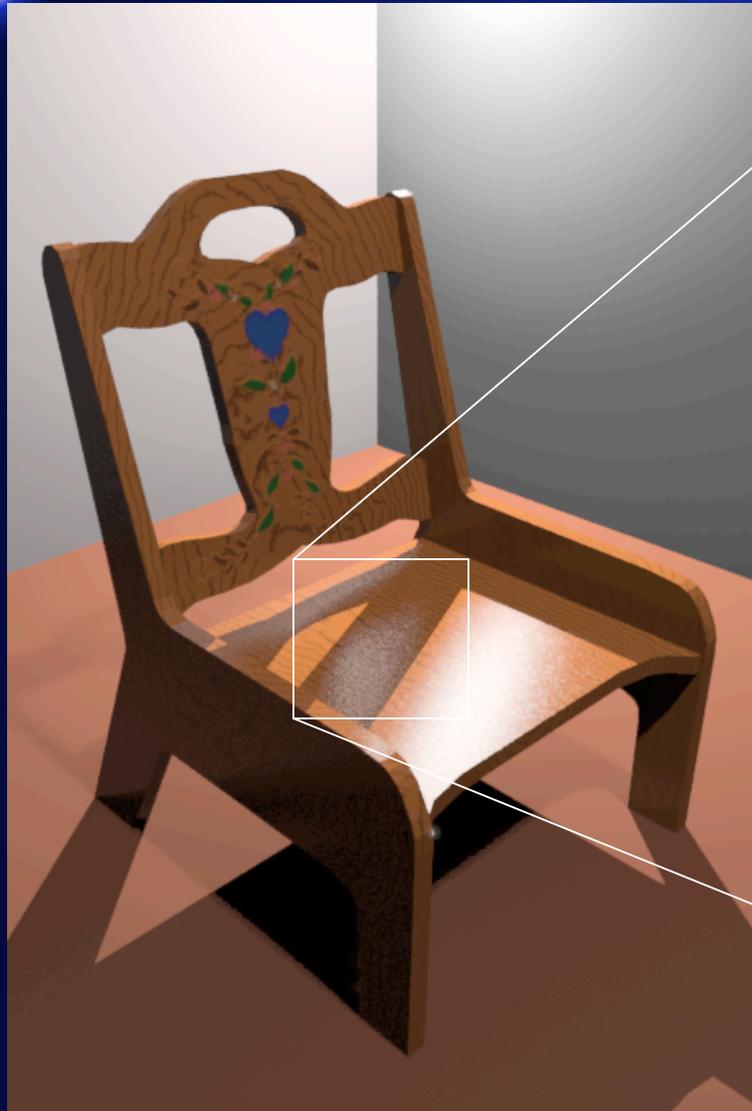
- Added pure Monte Carlo (`-u`) option
- Added depth-of-field (`-pd`) option
- Added “Russian roulette” ray termination
- Enhanced interreflection calculation
- **mksource** program to identify light sources
- **rtcontrib** program for contribution coeff.
- Created **meta2bmp** metafile converter
- Retired several unused programs

Quasi-Monte Carlo (-u-)



Familiar “brushed” appearance

Pure Monte Carlo (-u+)



Noisier but no discernable pattern

Depth-of-Field Sampling (-pd)

- Assigns aperture diameter in world units
- Also requires focus distance
 - Uses length of view direction (-vd) vector
 - Set by new rvu “focus” command
- More accurate than pdfblur with pinterp

Depth-of-Field Example

No blur



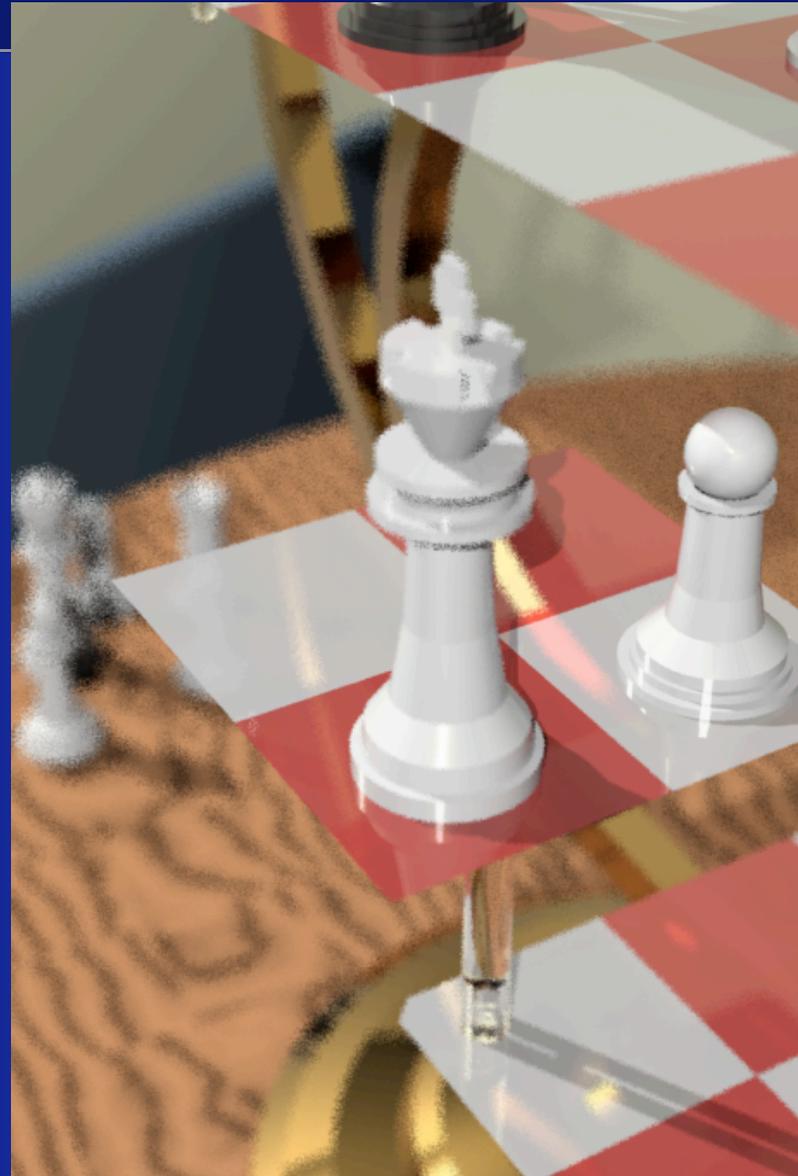
Depth-of-Field Example

**pdfblur with pinterp
(32 samples)**



Depth-of-Field Example

`rpict` with `-pd` option

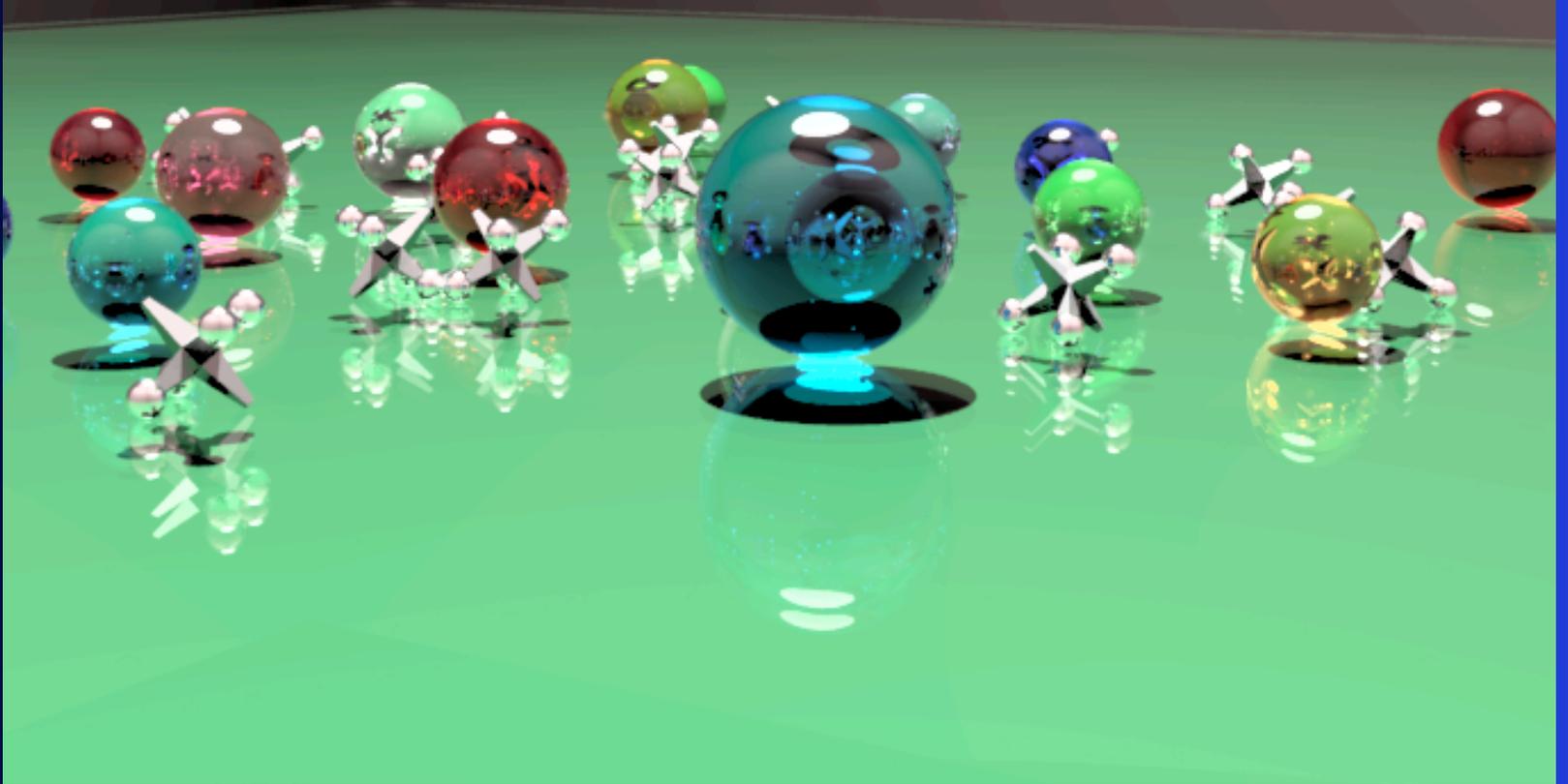


Russian Roulette Sampling

- Standard algorithm is strict termination below $-lw$ minimum weight limit
- Russian roulette continues trace with probability of (ray_weight/min_weight)
 - Continued ray weight is reset to min_weight
- Eliminates bias due to ray termination
- Enabled with negative value to $-lr$ option
- Made possible by code modifications needed for **rtcontrib**

Russian Roulette Example

Standard solution ($-lw \ .01 \ -lr \ 10, \text{Time} = 1.0$)



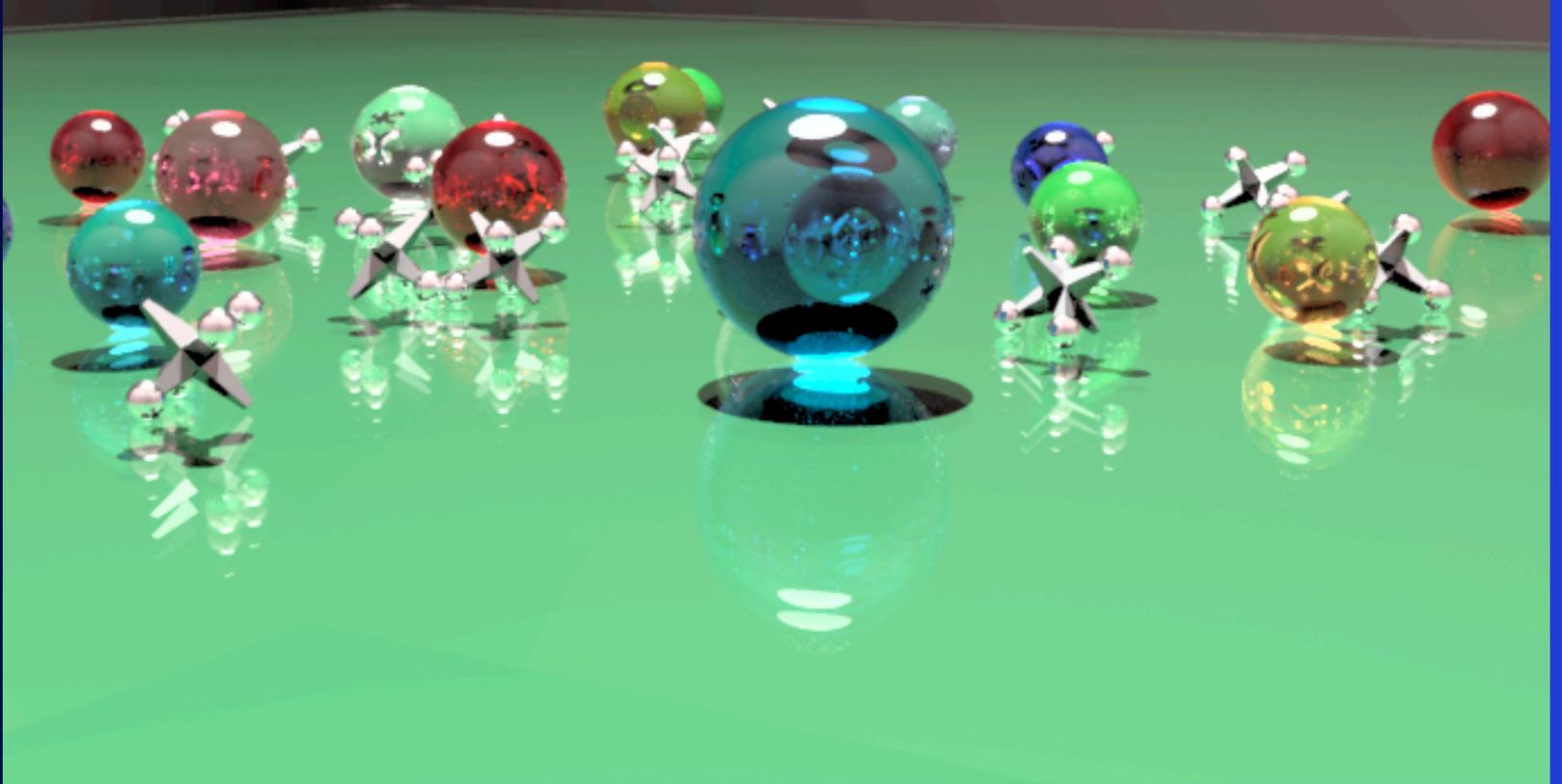
Russian Roulette Example

Ideal result ($-lw \ 1e-6 \ -lr \ 10$, Time = 4.3)



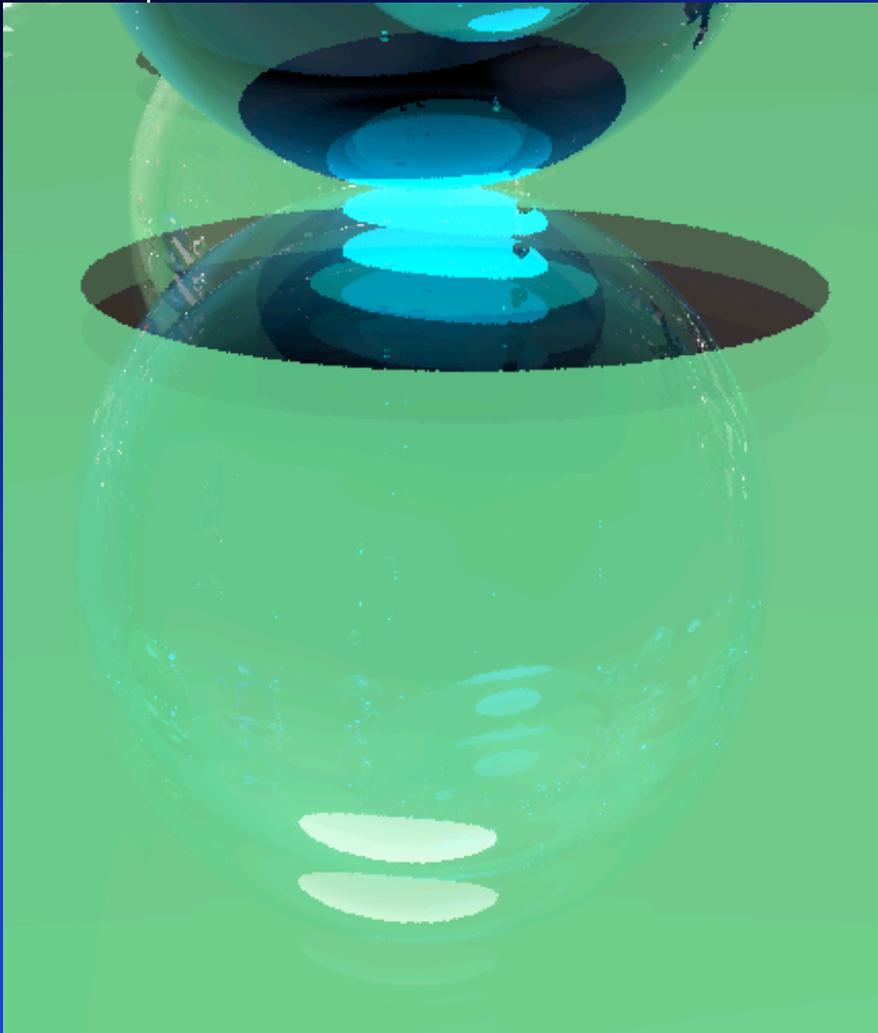
Russian Roulette Example

Russian roulette ($-1w \ .01 \ -1r \ -10$, Time = 1.3)



Russian Roulette Example

Ideal result



Standard Solution



mksource Program

- Identifies bright regions in HDR environment map (*light probe*)
- Superimposes distant *illum source*'s
- Reduces noise in resulting renderings
- **Only works with *glow source* inputs**

mksource Algorithm

- Sample all directions using geodesic mesh
 - Determine threshold from top 2%
 - Loop until no pixels over threshold:
 1. Identify brightest unclaimed pixel
 2. Grow source towards brightest unclaimed perimeter until:
 - a) Source exceeds maximum size, or
 - b) Perimeter values all below threshold, or
 - c) Source average drops below threshold
- + Methods to avoid over- and under-counting

Example mksource Results



Example mksource Results



Example mksource Results



rtcontrib Program

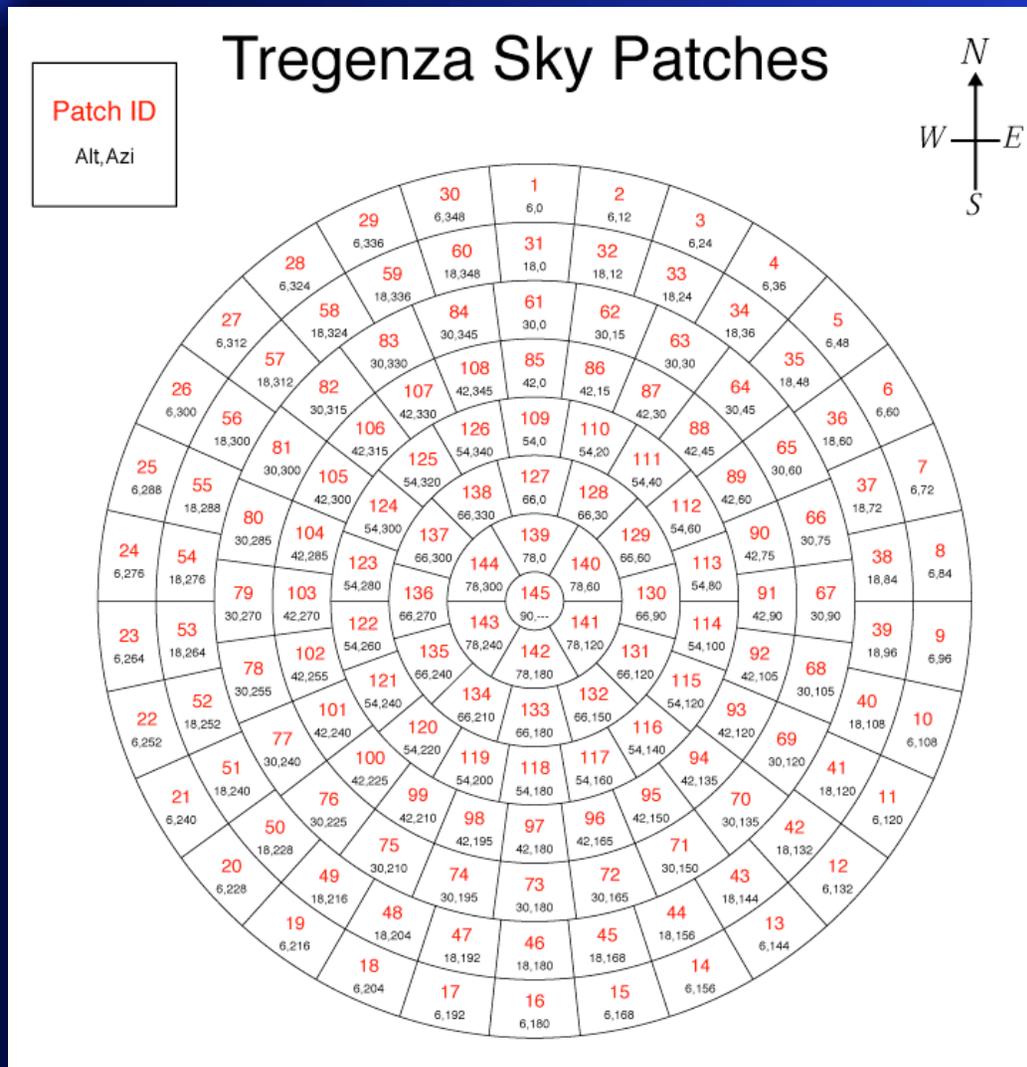
- Exploits new capability in **rtrace** for reporting ray contribution coefficients
- Contribution coefficients only useful deep in ray tree \Rightarrow 100+ million values per image
- **rtcontrib** sums & tabulates coefficients
 - Flexible control over where to gather rays
 - Output can be *Radiance* picture files, ASCII tables, or binary float's or double's
- **pcomb** useful in summing coeff. pictures

Simple rtcontrib Example

```
% rtcontrib @render.opt -I+  
-o c_%s.dat -m light1 -m light2  
scene.oct < test.dat
```

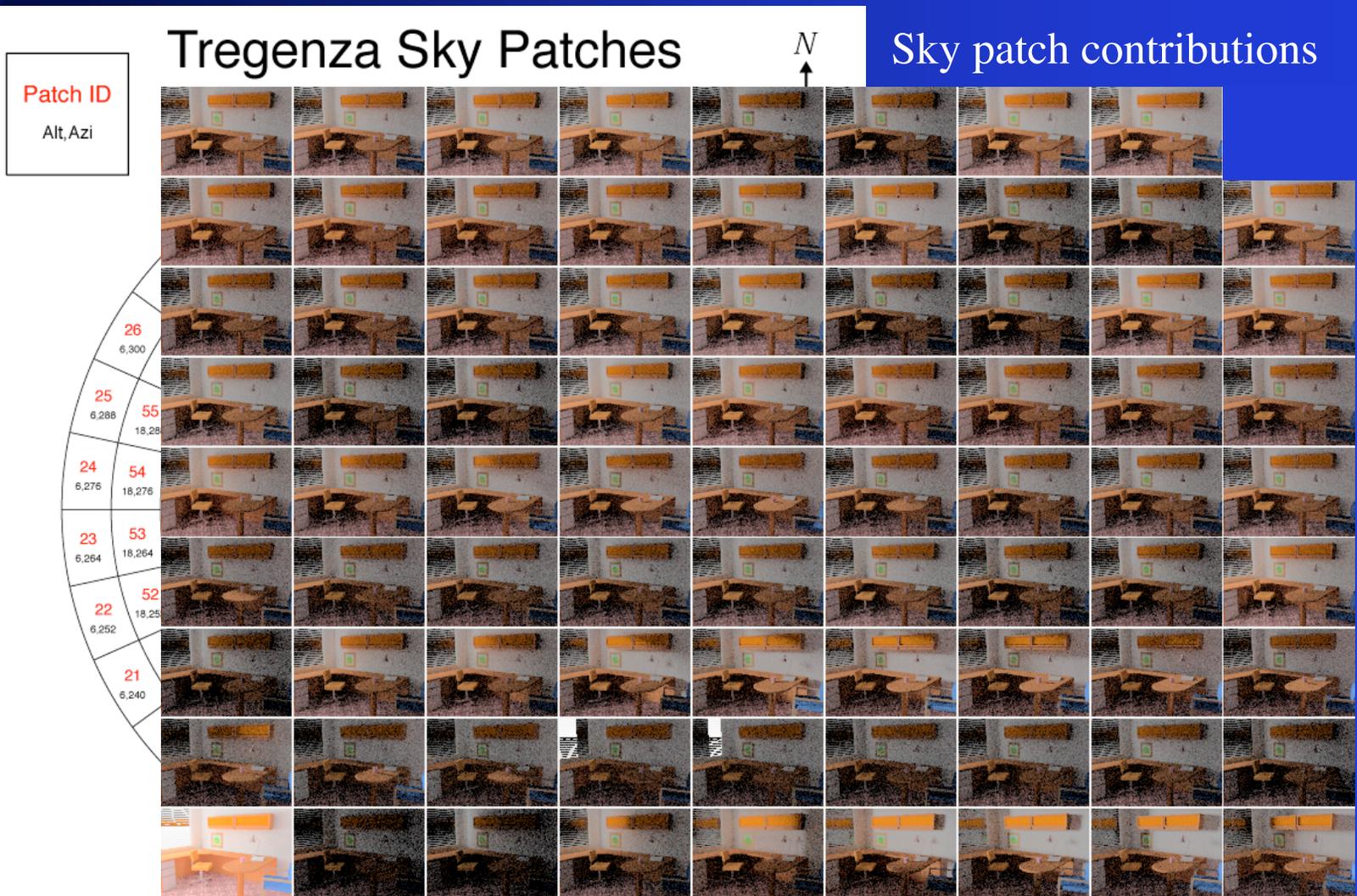
- Produces two files: c_light1.dat & c_light2.dat
- These contain irradiance coefficients for our two sources
- Linear combination of coefficients yields illuminance for any dimming settings
- Advantage: 1 calculation rather than 2

More Practical Example



Tregenza's Daylight Coefficient Method:
Sky is divided into roughly equal-sized patches, and the relative contribution of each patch is calculated for a given geometry

More Practical Example

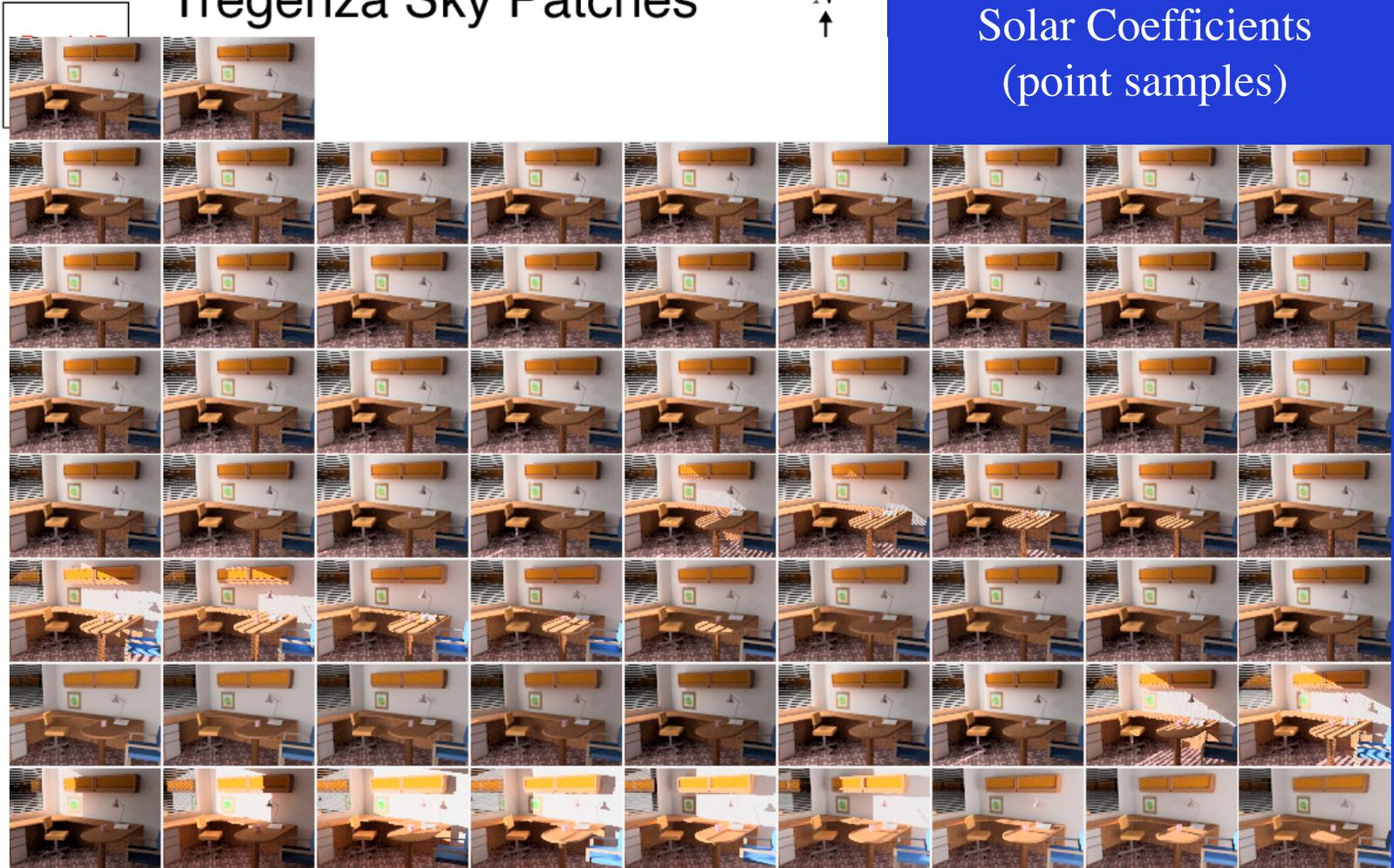


More Practical Example

Tregenza Sky Patches

N
↑

Solar Coefficients
(point samples)



Summed Result from pcomb



Further Possibilities

- Have not tried computing BRDF's and BTDF's, but sure it can be done
- Might be possible to extend **mkillum** to handle curved, specular devices
- More details on how to use **rtcontrib** with a more extensive example in afternoon talk